

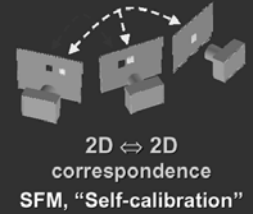
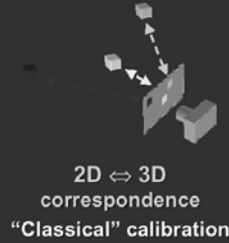
## Camera Calibration

EN193s08 3D Photography  
Brown Fall 2003  
Gabriel Taubin

## Geometric Camera Calibration

### Augmented pin-hole camera model

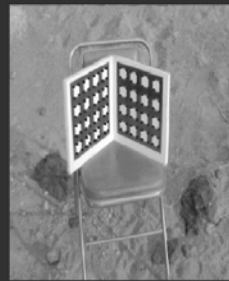
- Focal point, orientation
- Focal length, aspect ratio, center, lens distortion



## Camera Calibration

- **Geometry**
  - Where is the camera located ?
  - How is the camera oriented ?
  - What are the internal parameters ?
- **Radiometry**
  - What is the relation from incident light intensity to pixel values ?
- Parameters are estimated by observing objects of known dimensions and measuring their projections on the image plane

## Calibration Patterns



Calibration grid  
Z. Zhang, Microsoft Research



Chromaglyphs  
Bruce Culbertson, HP-labs

## Calibration From 2D Motion



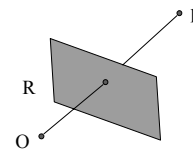
### Structure From Motion (SFM)

- Track points over a sequence of images
- Solve for 3D positions and camera positions
- Calibrate internal parameters beforehand

### Self-Calibration

- Solve for internal *and* external parameters
- E.g., [Pollefeys, 98]

## Extrinsic Parameters



$$\begin{bmatrix} x \\ y \\ z \end{bmatrix} = \begin{bmatrix} R_{11} & R_{12} & R_{13} \\ R_{21} & R_{22} & R_{23} \\ R_{31} & R_{32} & R_{33} \end{bmatrix} \begin{bmatrix} p_x - O_x \\ p_y - O_y \\ p_z - O_z \end{bmatrix}$$

3 DOF Rotation ( $R=R(v)$ ) + 3 DOF Position ( $O$ )

### Intrinsic Parameters Without lens distortion

$$u = f_u (x/z) + c_u$$

$$v = f_v (y/z) + c_v$$

2 DOF Focal Lengths (f)  
2 DOF Center of Projection (O)

### Intrinsic Parameters Without lens distortion

$$s \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \begin{bmatrix} f_u & 0 & c_u \\ 0 & f_v & c_v \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \end{bmatrix}$$

2 DOF Focal Lengths (f)  
2 DOF Center of Projection (O)

### Projection Matrix

$$s \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \underbrace{\begin{bmatrix} f_u & 0 & c_u \\ 0 & f_v & c_v \\ 0 & 0 & 1 \end{bmatrix}}_{P_I} \underbrace{\begin{bmatrix} R_{11} & R_{12} & R_{13} \\ R_{21} & R_{22} & R_{23} \\ R_{31} & R_{32} & R_{33} \end{bmatrix}}_{P_E} \begin{bmatrix} p_x - O_x \\ p_y - O_y \\ p_z - O_z \end{bmatrix}$$

### Linear Geometric Calibration

$$s \begin{bmatrix} u \\ v \\ 1 \end{bmatrix} = \underbrace{\begin{bmatrix} P_{11} & P_{12} & P_{13} & P_{14} \\ P_{21} & P_{22} & P_{23} & P_{24} \\ P_{31} & P_{32} & P_{33} & P_{34} \end{bmatrix}}_P \begin{bmatrix} p_x \\ p_y \\ p_z \\ 1 \end{bmatrix}$$

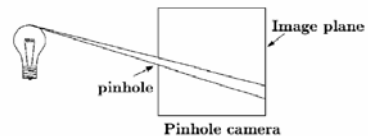
- Know 3D coords and 2D coords
- 11 unknowns (up to scale)
- 2 equations per point (eliminate s)
- 6 points are sufficient

### Nonlinear Methods

- Problems with Linear Method
  - Too many free parameters
  - Doesn't model lens distortion
- Nonlinear Methods [Tsai, 1985]
  - Parameterize matrix P as a function of
    - Rotation vector v
    - Position vector O
    - Focal length vector f
    - Center of projection c
    - Lens distortion parameters
  - Use non-linear Least-Squares Solver

### The pinhole camera

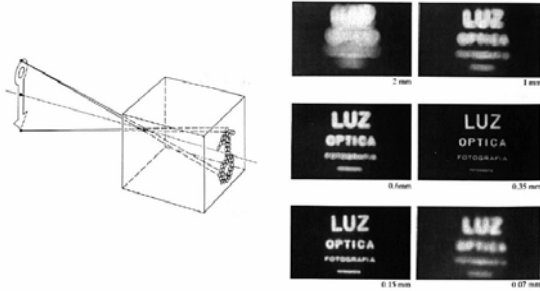
The first camera - "camera obscura" - known to Aristotle.



**Small aperture = high fidelity**  
*but requires long exposure or bright illumination*

## Pinhole camera

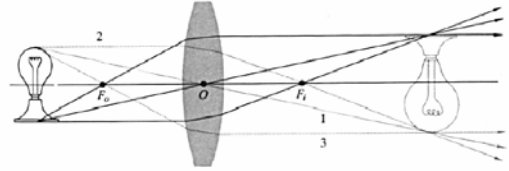
If aperture is too small, then diffraction causes blur.



[Figure from Hecht87]

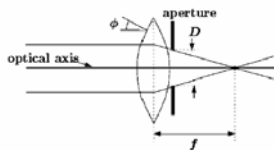
## Lenses

Lenses focus a bundle of rays to one point.  
=> can have larger aperture.



[Figure from Hecht87]

## Lenses



A lens images a bundle of parallel rays to a focal point at a distance,  $f$ , beyond the plane of the lens.  
Note:  $f$  is a function of the index of refraction of the lens.

An aperture of diameter,  $D$ , restricts the extent of the bundle of refracted rays.

## Lenses

For economical manufacture, lens surfaces are usually spherical.

A spherical lens behaves ideally if  $\phi$  is small:

$$\sin \phi = \phi - \frac{\phi^3}{3!} + \frac{\phi^5}{5!} - \dots \approx \phi$$

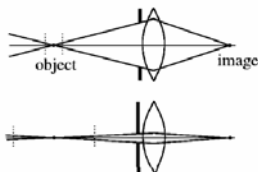
The angle restriction means we consider rays near the optical axis -- "paraxial rays."

## Depth of field

Lens systems do have some limitations.

First, points that are not in the object plane will appear out of focus.

The *depth of field* is a measure of how far from the object plane points can be before appearing "too blurry."



## Monochromatic aberrations

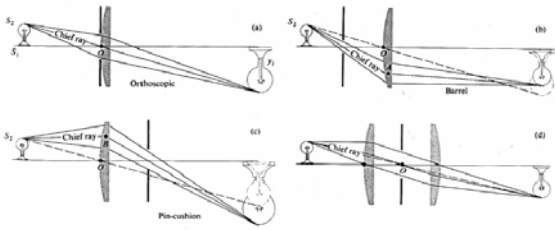
Allowing for the next higher terms in the  $\sin \phi$  approximation:

$$\sin \phi = \phi - \frac{\phi^3}{3!} + \frac{\phi^5}{5!} - \dots \approx \phi - \frac{\phi^3}{3!}$$

...we arrive at the third order theory. Deviations from ideal optics are called the *primary* or *Seidel aberrations*:

- Spherical aberration
- Coma
- Astigmatism
- Petzval curvature
- Distortion

## Distortion



[Figures from Hecht87]

## Chromatic aberration

**Cause:**

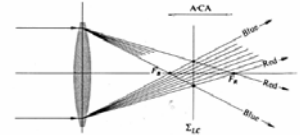
Index of refraction varies with wavelength.

**Effect:**

Focus shifts with color, colored fringes on highlights

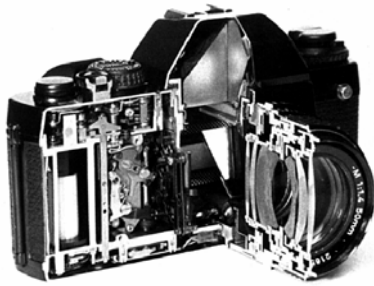
**Ways of improving:**

Achromatic designs



[Figure from Hecht87]

## The art of optical design...



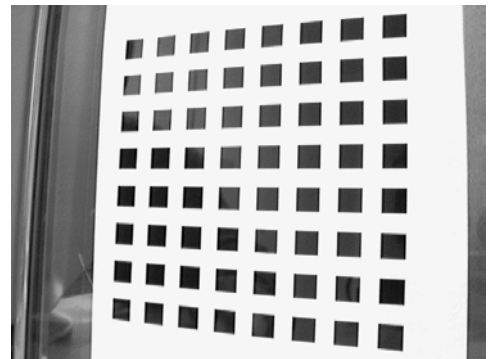
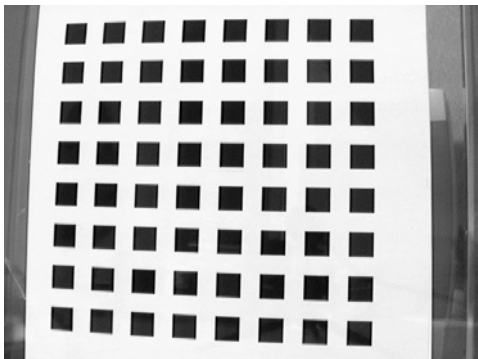
[Figure from Goldberg92]

## Intrinsic Parameters With Lens Distortion

$$u = f_u \phi_u(x/z, y/z) + c_u$$

$$v = f_v \phi_v(x/z, y/z) + c_v$$

2 DOF Focal Lengths ( $f$ )  
 2 DOF Center of Projection ( $O$ )  
 ? DOF Lens Distortion Function ( $\phi$ )



### Intrinsic Parameters With Lens Distortion

$$u' = f_u (x/z) + c_u$$

$$v' = f_v (y/z) + c_v$$

$$u = \rho_u (u', v')$$

$$v = \rho_v (u', v')$$

If lens distortion is moved to the last step,  
It can be estimated independently as the first step,  
And images can be warped after capture to remove distortion

### Why is lens distortion independent ?

- Projection is now
  - a projective transformation from 3D to 2D
  - Followed by a 2D to 2D warping (lens distortion)
- Warping can be estimated and corrected by looking at the image of a planar pattern or lines because
- Homography from plane containing the pattern to image plane preserves straight lines
- So, the problem is to find the warping parameters that straighten all the lines as much as possible